

# Measuring large-scale UV background inhomogeneities with the Hell forest and metals

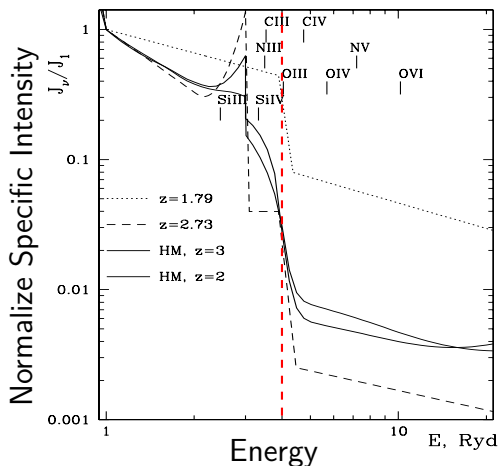
Sean Morrison<sup>1</sup>  
Mat Pieri<sup>1</sup>, David Syphers,<sup>2</sup> & Tae-Sun Kim<sup>3</sup>

<sup>1</sup>Laboratoire d'Astrophysique de Marseille, Université d'Aix-Marseille,  
<sup>2</sup>Eastern Washington University,  
<sup>3</sup>University of Wisconsin-Madison

Intergalactic Interconnections 12 July 2018

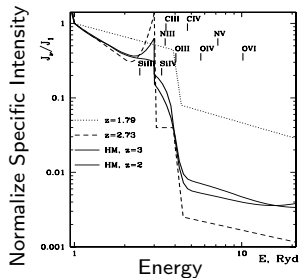
# UV Background Shape & Intensity

- H I and He II reionization
- Quasar population, galaxy population
- Traced using various ionization species



Agafonova et al. (2007)

# Quasar/Transverse Proximity Effects

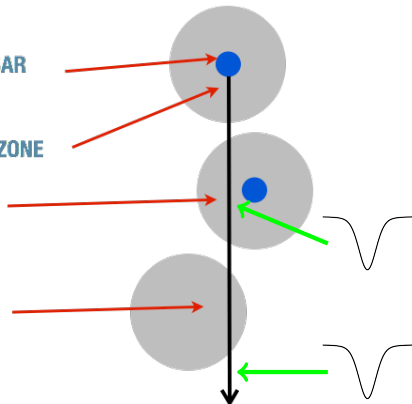


**SOURCE QUASAR**

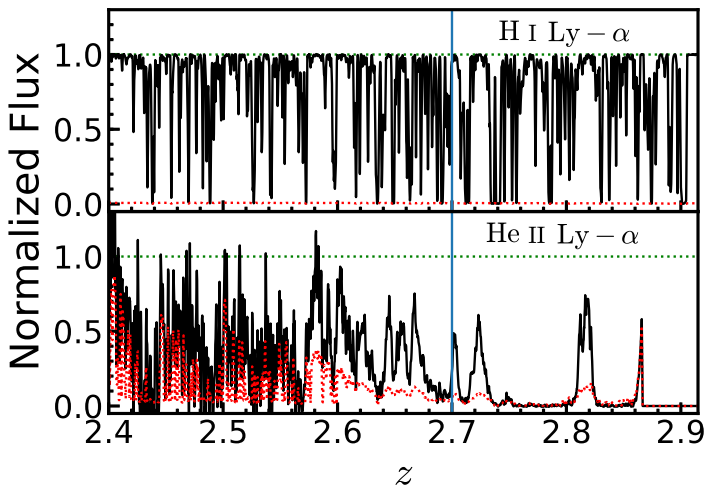
**PROXIMITY ZONE**

**TRANSVERSE PROXIMITY ZONE**

**TRANSVERSE PROXIMITY ZONE OF UNOBSERVED AGN**



## Gunn-Peterson Trough in HE2347-4342



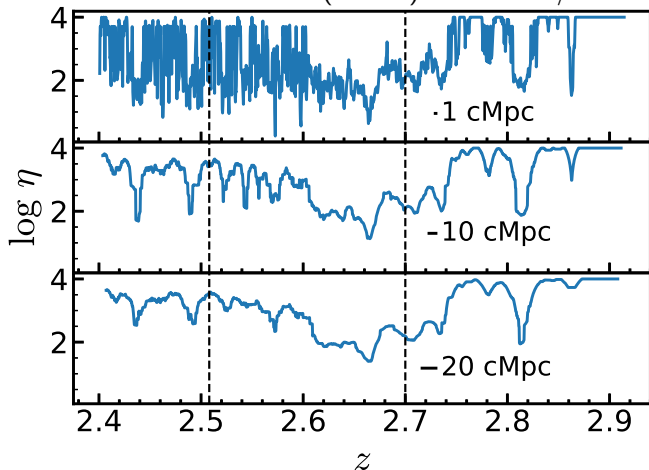
# $\eta$ : The Ratio of He II to H I

$$\eta = \frac{N_{\text{He II}}}{N_{\text{H I}}} \approx \frac{4\tau_{\text{He II}}}{\tau_{\text{H I}}}$$

- Density independent
- Large  $\eta$ : soft radiation
- Small  $\eta$ : hard radiation
- $\eta$  sensitive on
  - Mpc scales** to the locations of transverse quasars
  - smaller scales** to thermal broadening, galactic outflows, and proximity to local galaxies

Measuring  $\eta$  on Large Scales

- Measuring He Ly- $\alpha$  to H Ly- $\alpha$  ( $\eta$ ) ratio smoothed on  $> 1$  Mpc scales.
- Data from Hubble COS (for He) and UVES/HIRES (H)



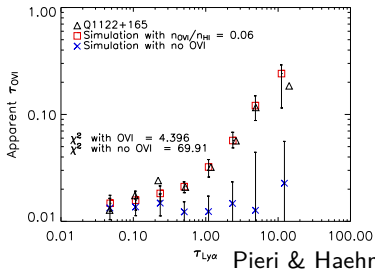
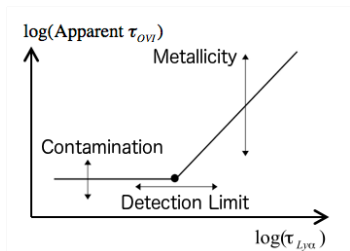
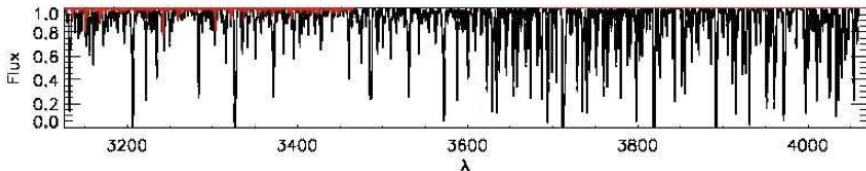
HE 2367-4342

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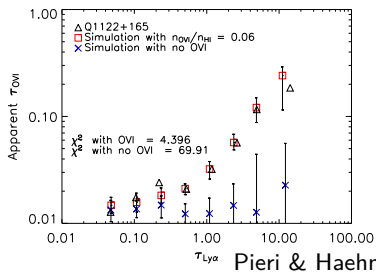
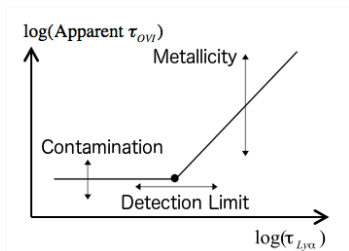
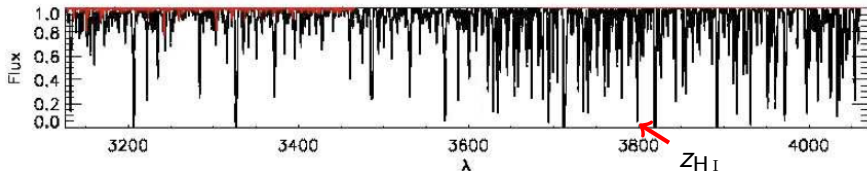
# Measuring Metals in the IGM

- Pixel Optical Depth Techniques (Songaila et al. 1995; Schaye et al. 2003)
- A example (simulated) Ly- $\alpha$  forest in a quasar spectrum  
with an OVI forest (5 times ionized oxygen)



# Measuring Metals in the IGM

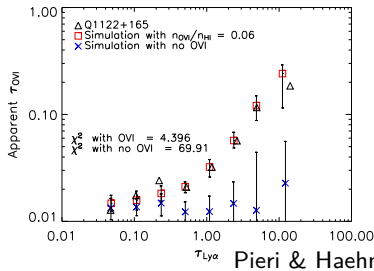
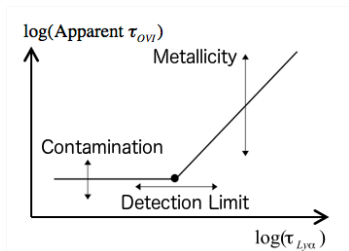
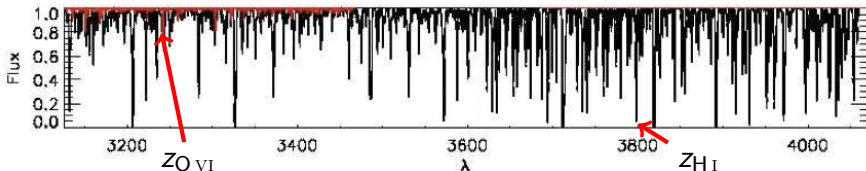
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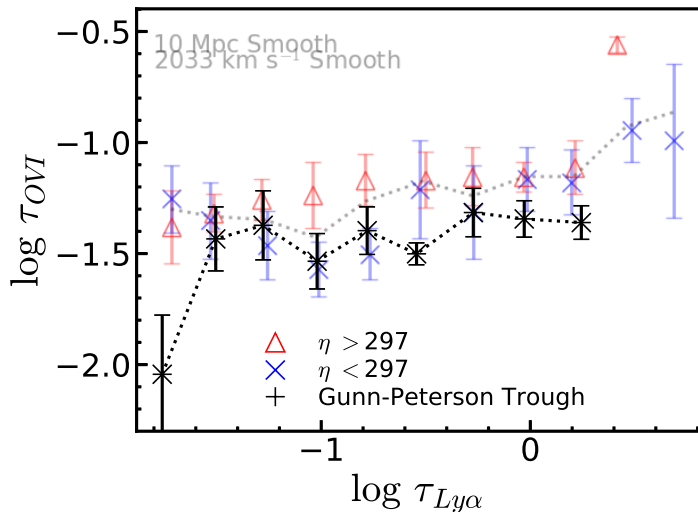


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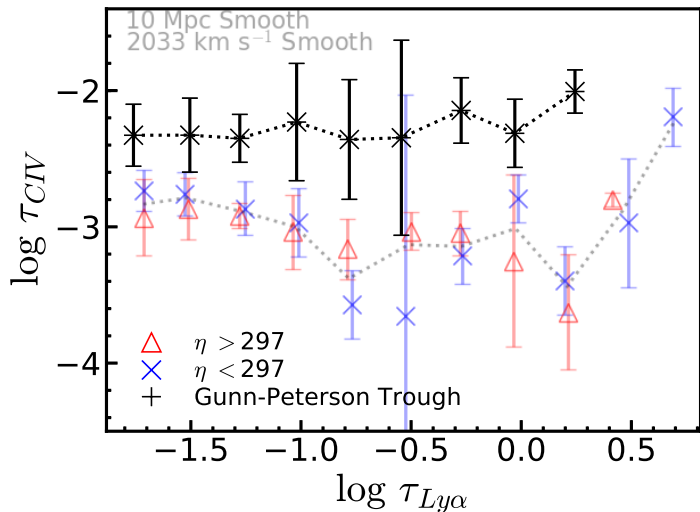
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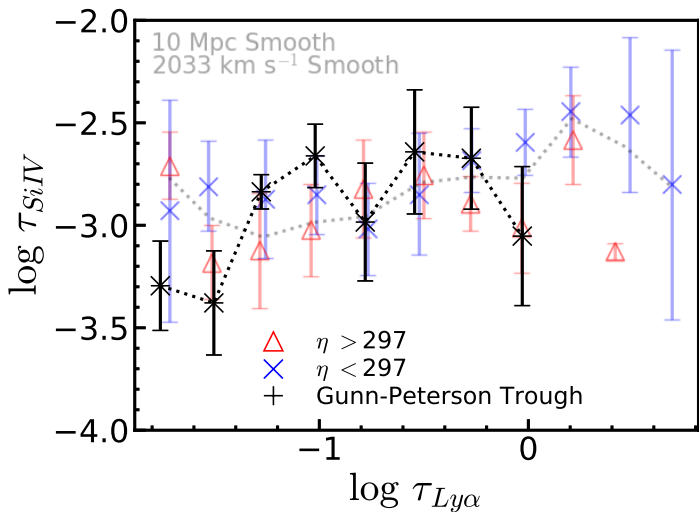
## POD in the Gunn-Peterson Trough: O VI



## POD in the Gunn-Peterson Trough: C IV

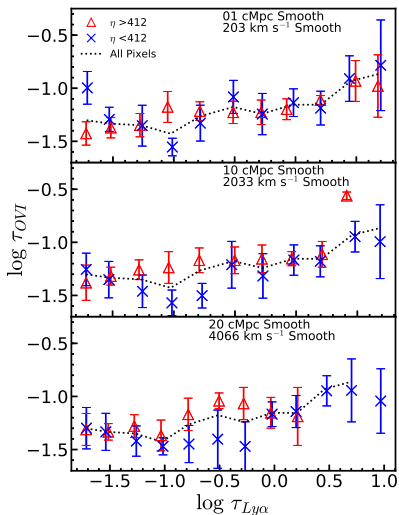


## POD in the Gunn-Peterson Trough: Si IV

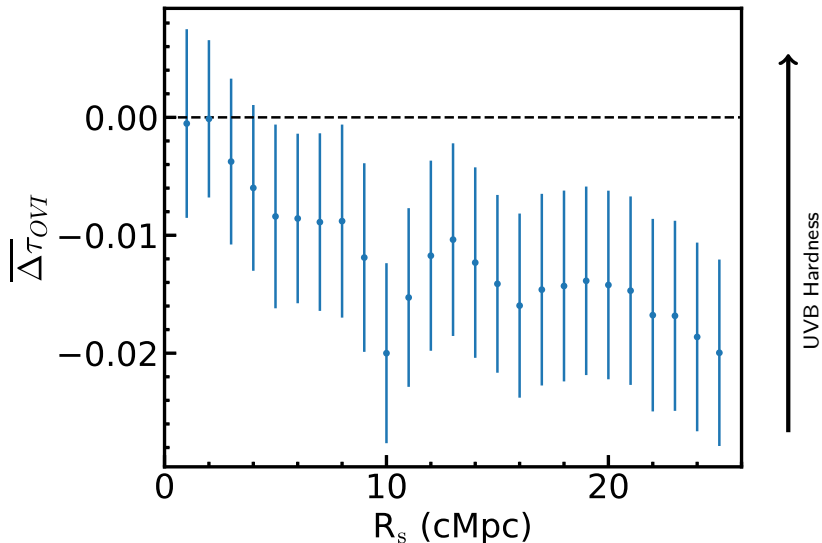


## UVB Soft/Hard Split: Raw POD

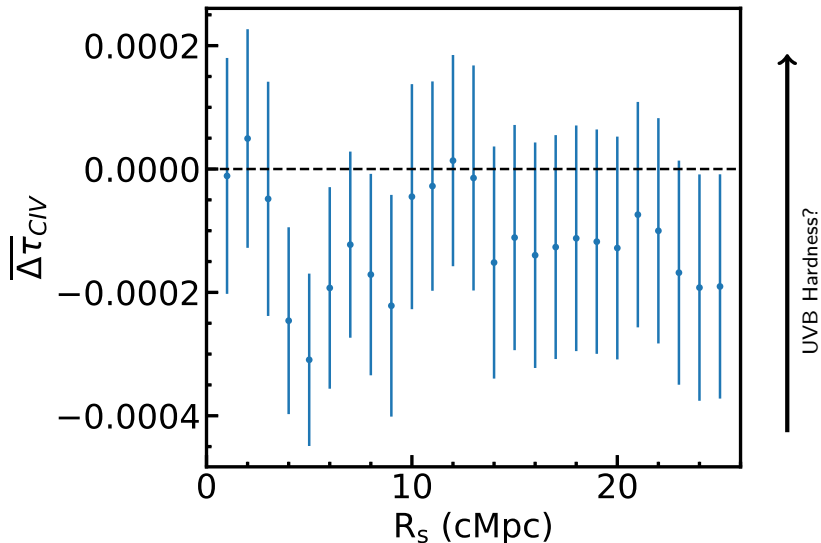
Statistical measurement of O VI absorption cut by high/low  $\eta$ .



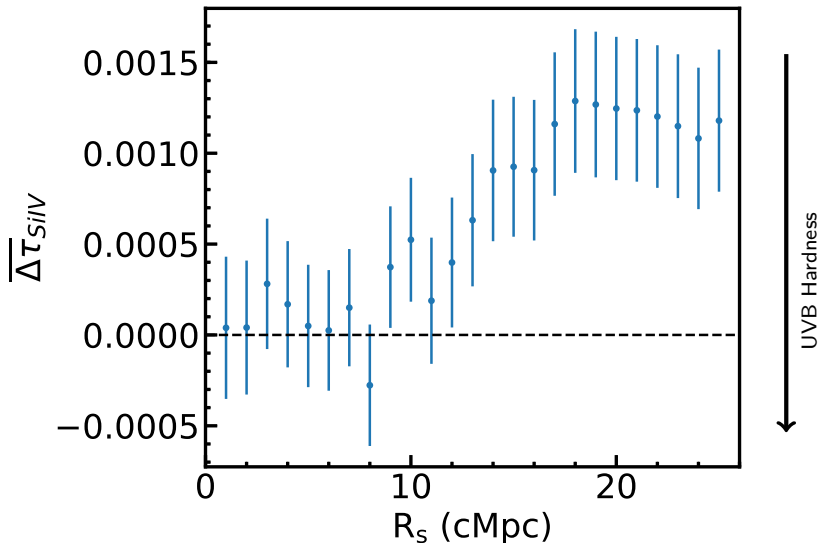
## Soft/Hard Split as a Function of Scale: O VI



## Soft/Hard Split as a Function of Scale: C IV

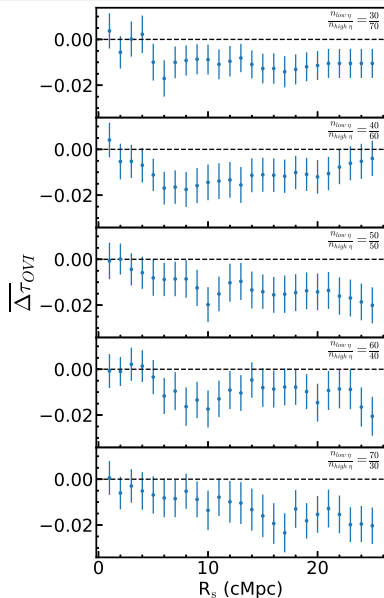


## Soft/Hard Split as a Function of Scale: Si IV





## Variable Soft/Hard Splits in O VI



- Difference in Metal (O VI & C IV) Column Densities in and out of Gunn-Peterson Trough
- Seems to support claim that  $\eta$  actually probes large scale effects rather than just noise (which has been source of controversy)
- (Potential) future work:
  - Comparison with simulations
  - Better S/N UV data for additional He II Quasars
  - Quasars positions in large surveys