

The physical conditions of gas flows observed with MUSE and ALMA

Celine Peroux

Martin Zwaan, Anne Klitsch, Ramona

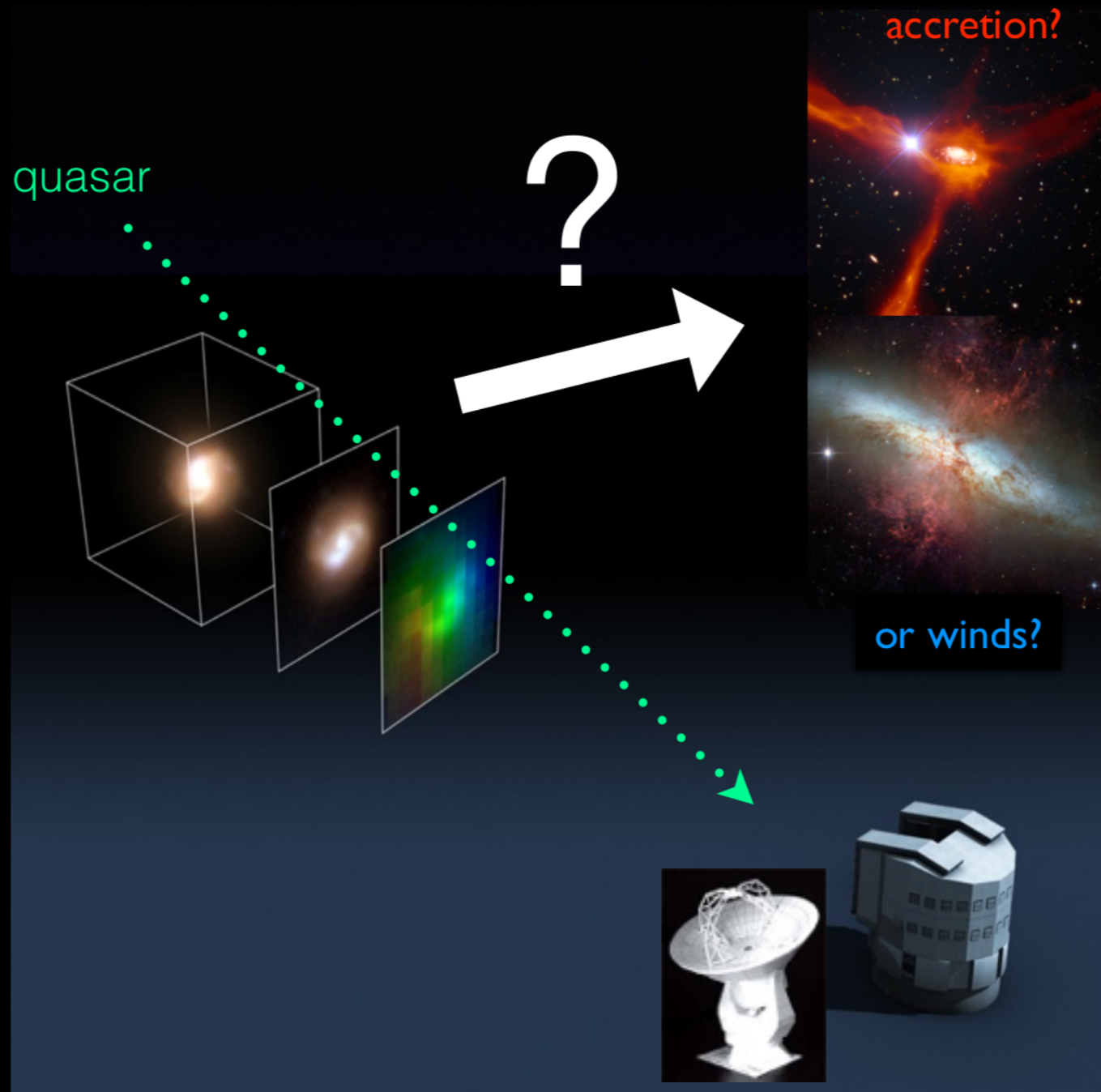
Augustin, Aleksandra Hamanowicz, Hadi Rahmani, Max Pettini,
Varsha Kulkarni, Lorrie Straka, Andy Biggs, Don York & Bruno Milliard



Questions to Address

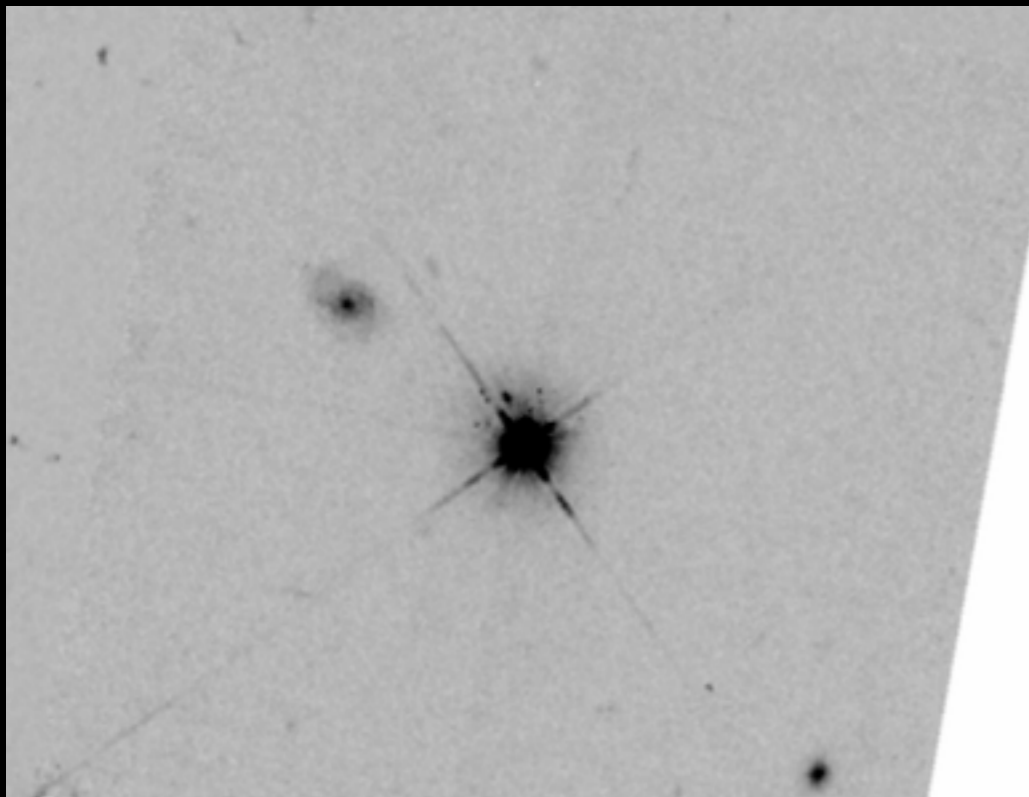
- 1- How to probe galactic gas flows?
- 2- How to characterise the multi-phase CGM?
- 3- On which scales are metals mixed?

3D cubes: a powerful tool to connect gas & stars



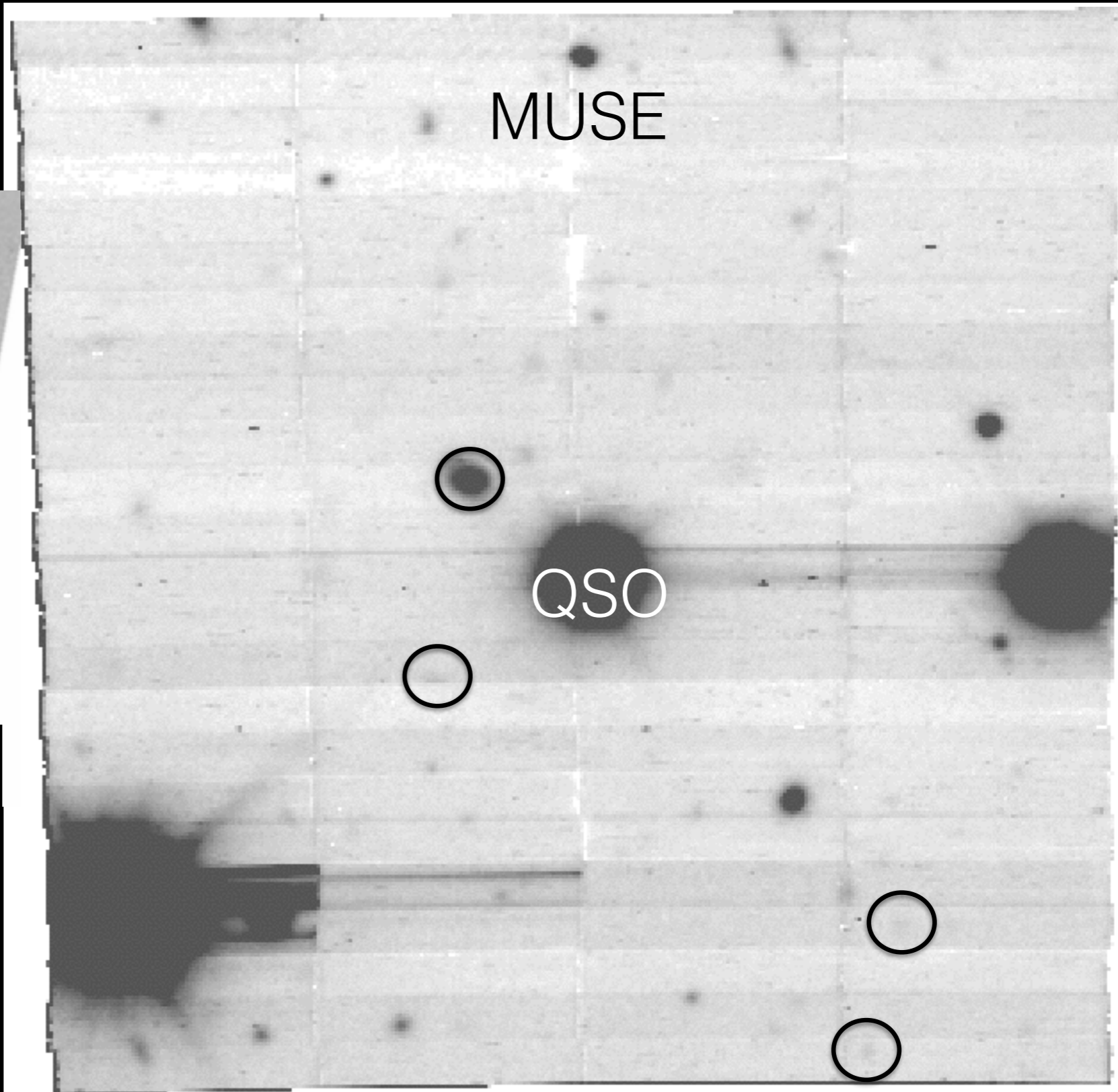
Gas Flows probe by Kinematics

HST



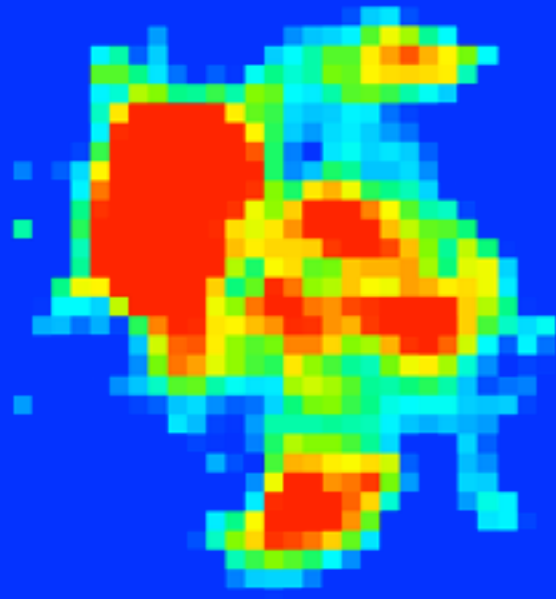
$\log N(\text{HI})=19.5$

MUSE



[OIII] 5008 Obs Flux

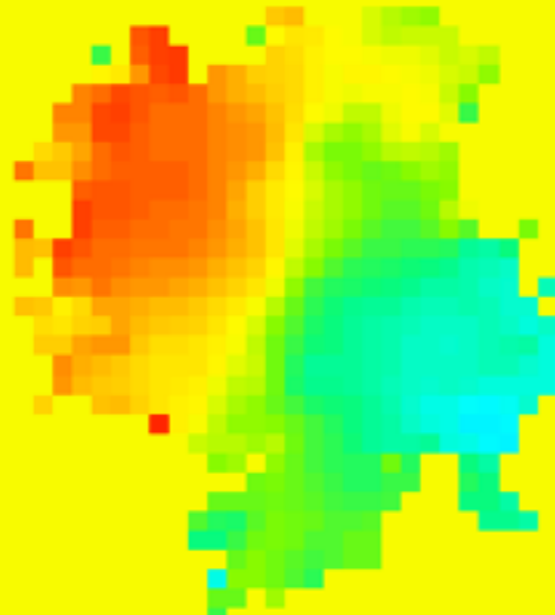
2.5



1.0×10^{-19}

Obs Velocity

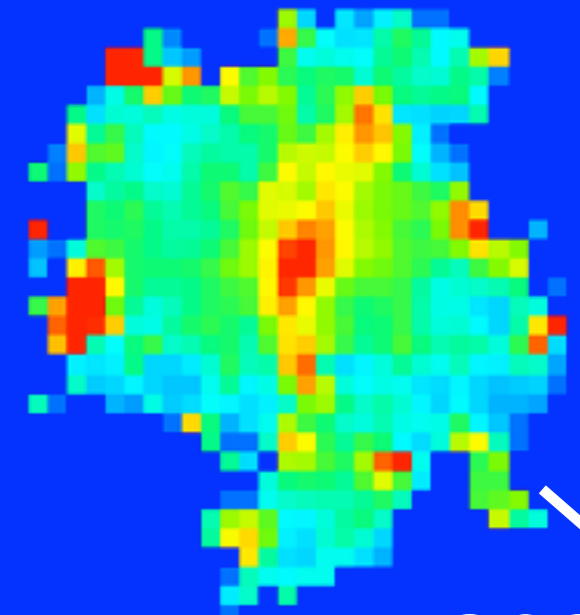
150



-150

Vel Dispersion

90



60

40

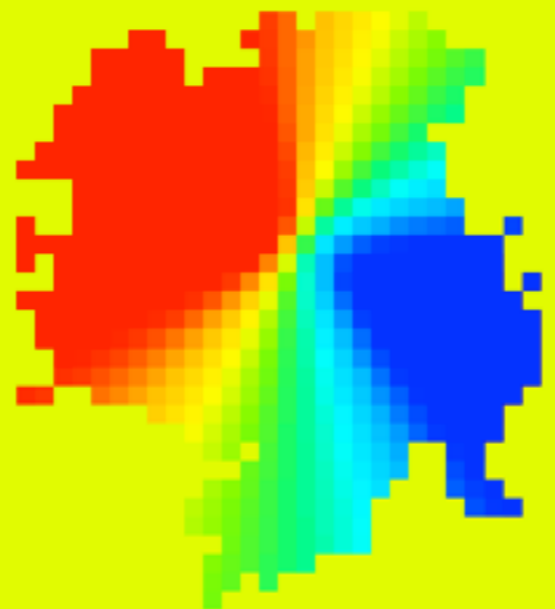
km/s

QSO

• $M_{\text{dyn}} = 7 \times 10^{10} M_{\text{sun}}$, $M_{\text{halo}} = 3 \times 10^{12} M_{\text{sun}}$

Model velocity

150

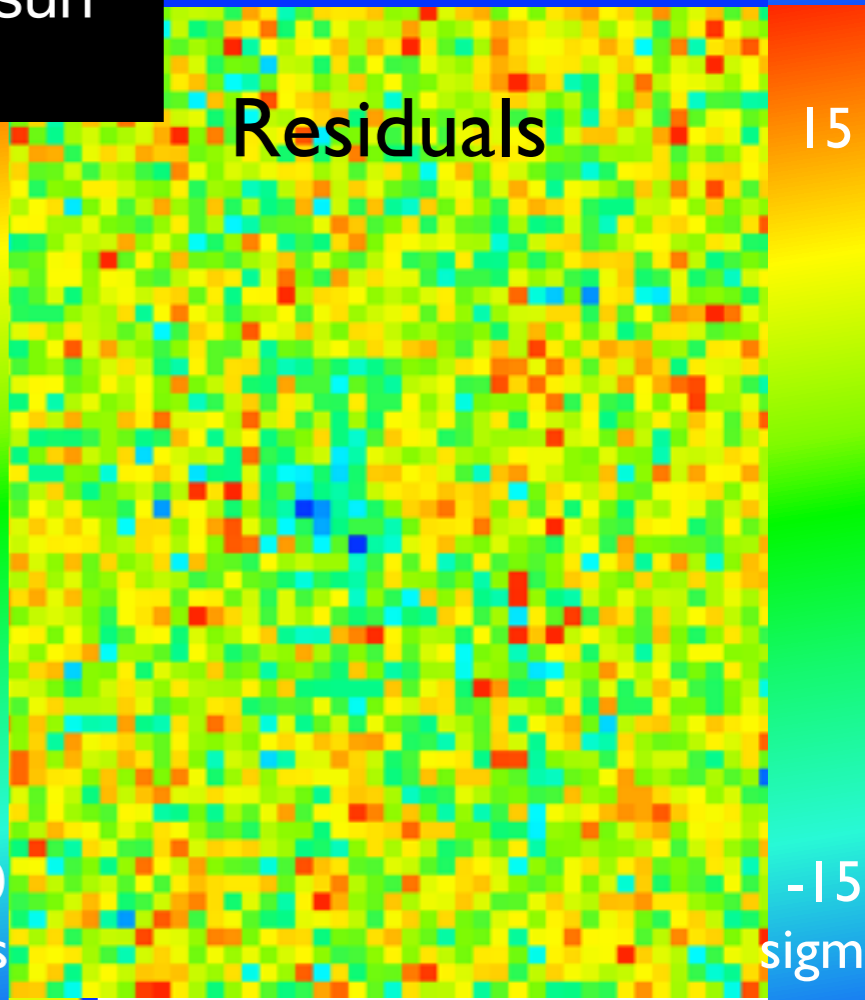


-150

km/s

Residuals

15

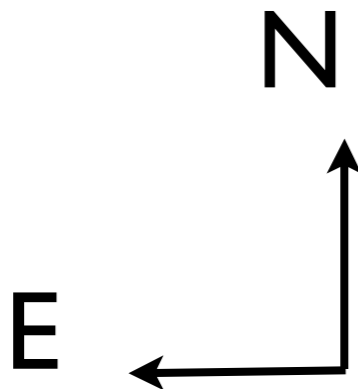


-15

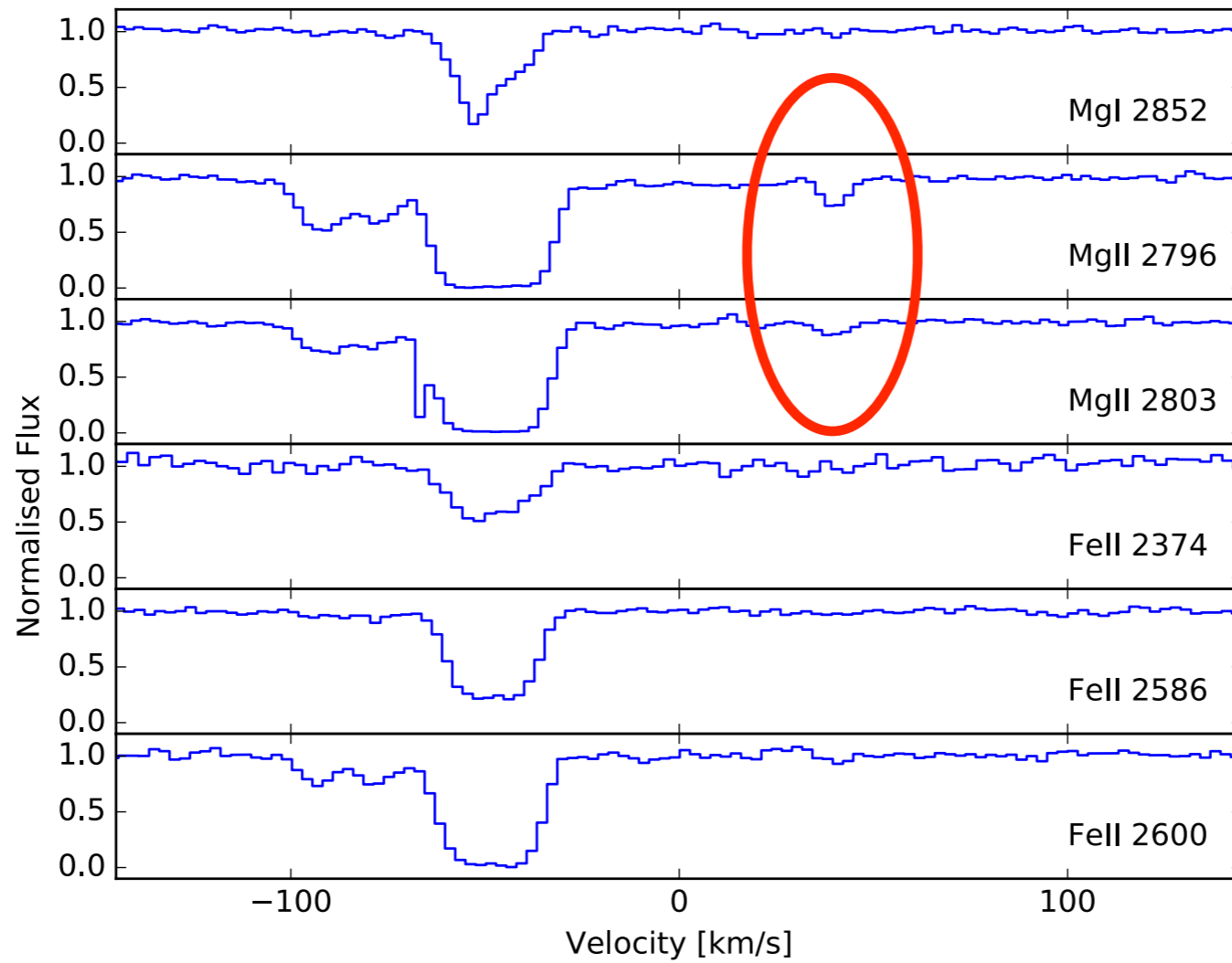
sigma



Galaxy "a"

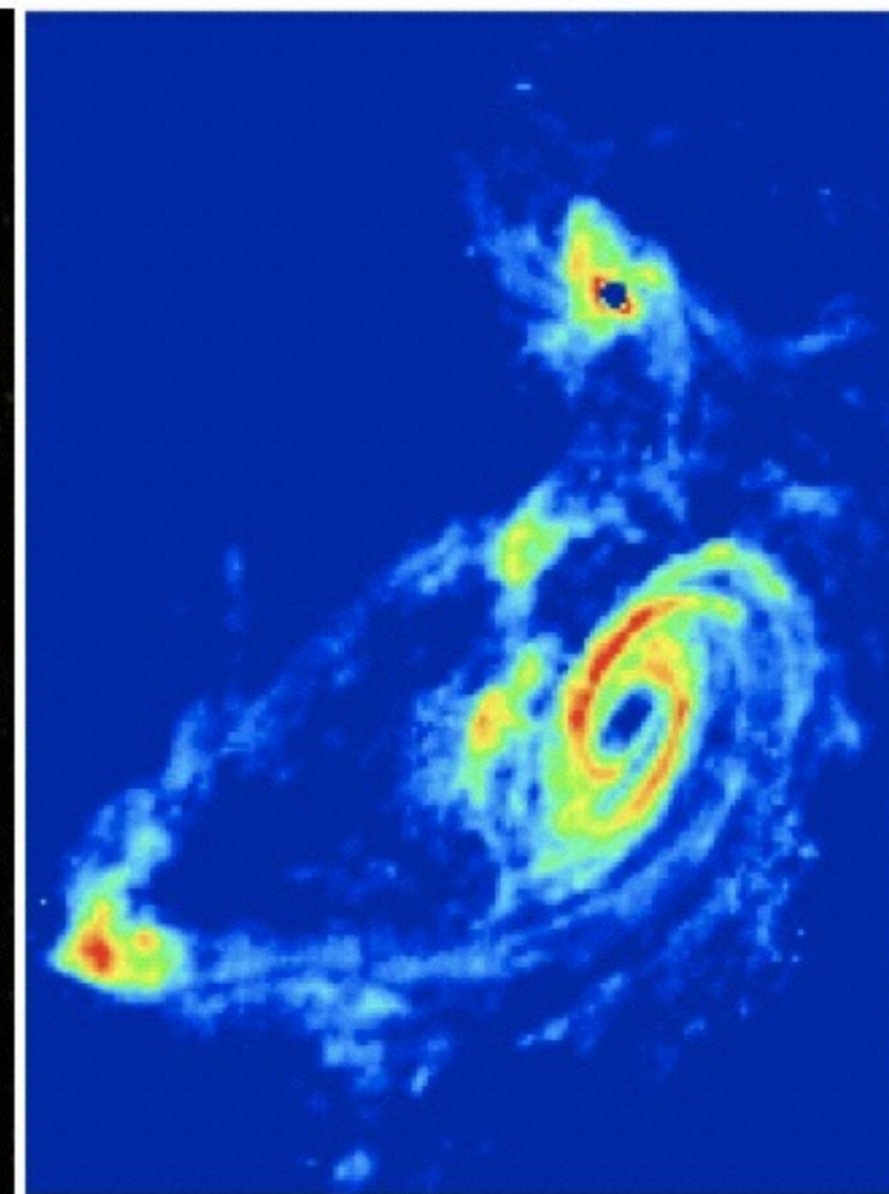
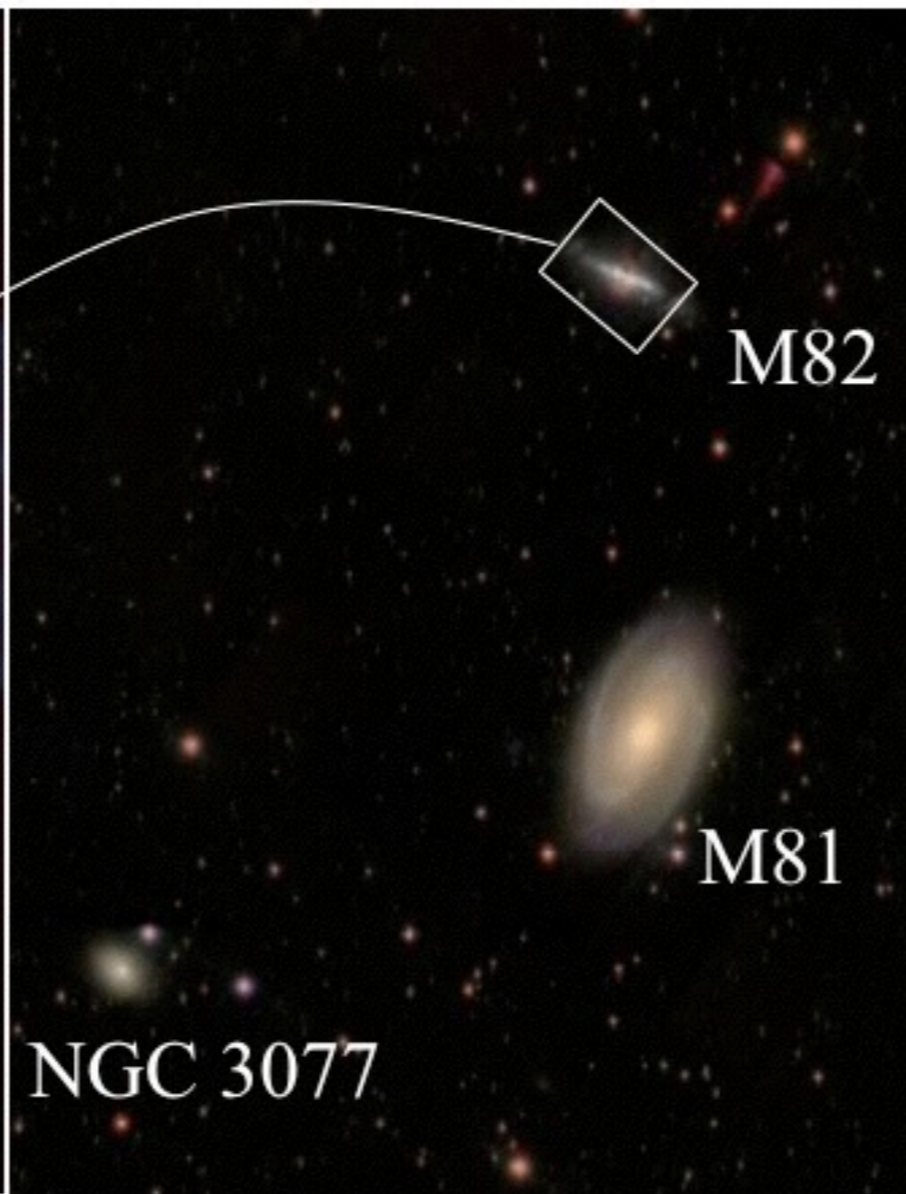
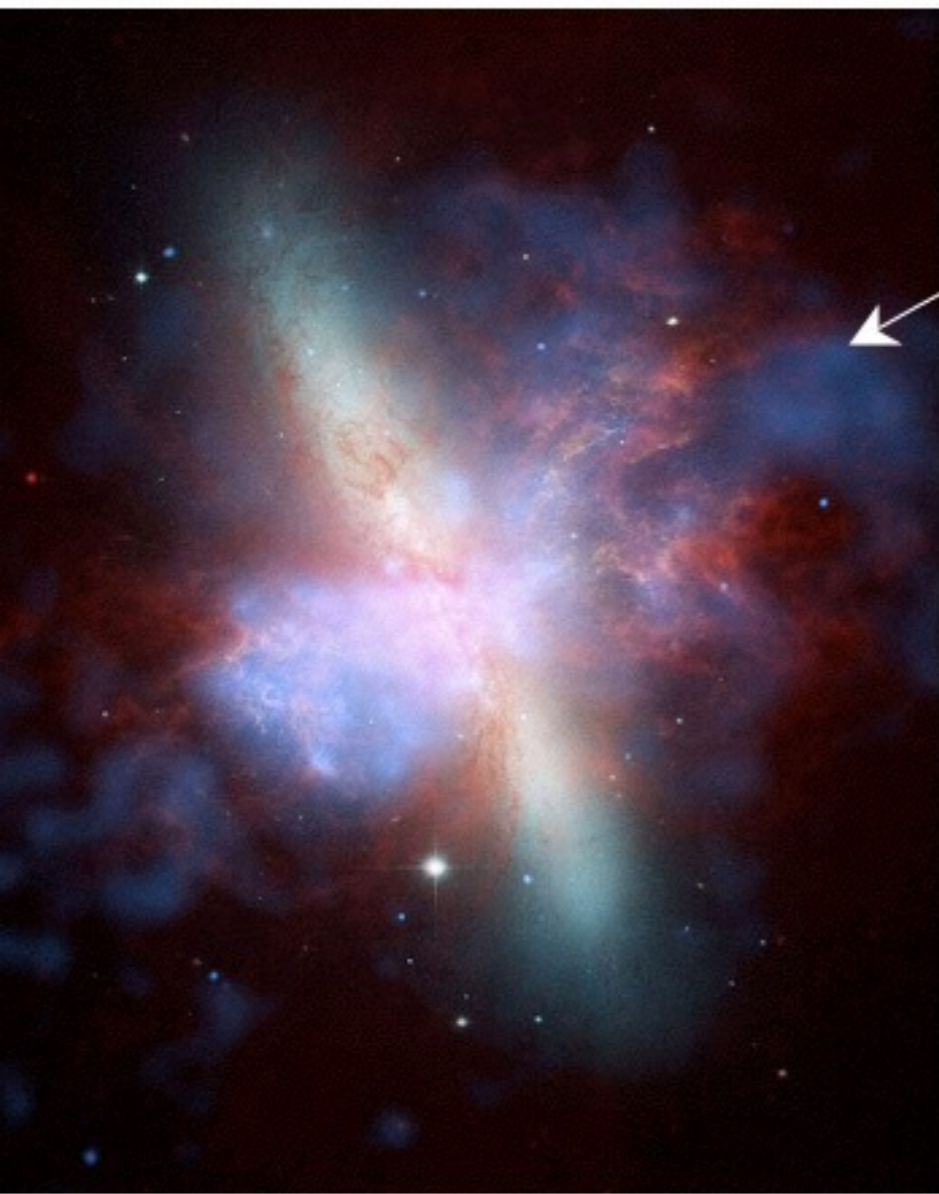


What is this Gas?



CP+17

Low-redshift Analog



Galactic Wind
(M82)

Starlight
(optical)

HI Gas
(radio)

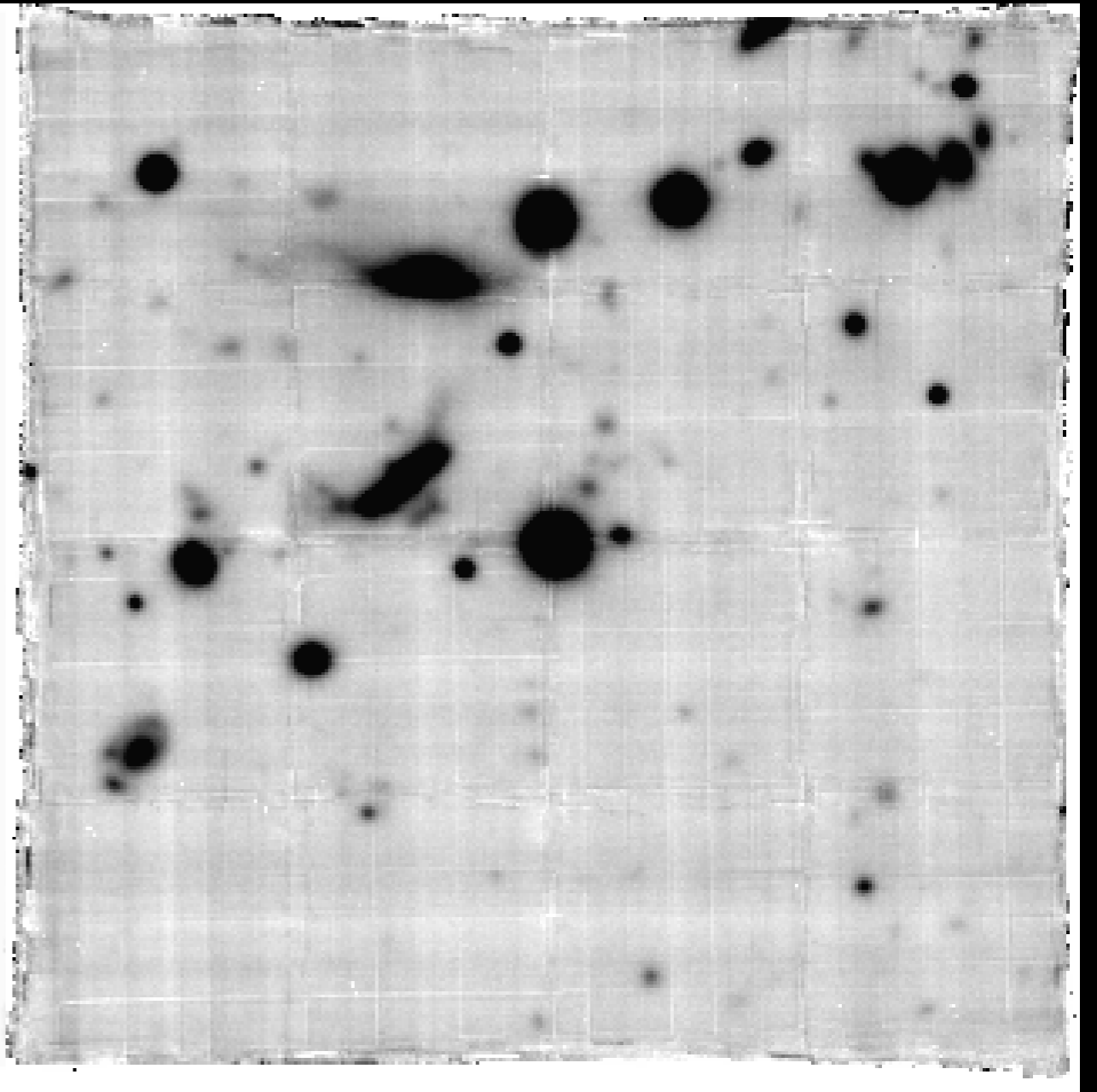
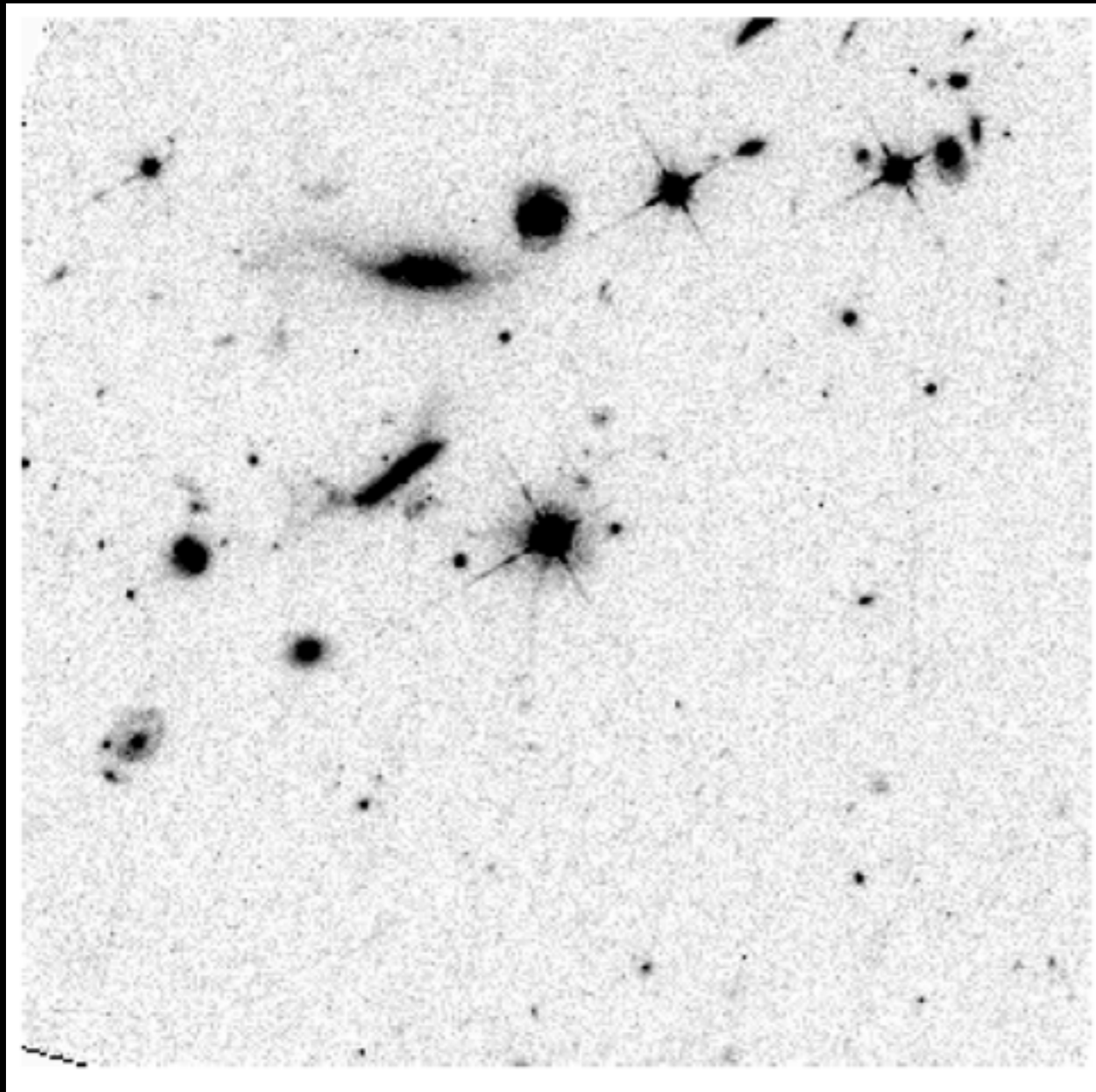
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Multi-Phase CGM

HST

MUSE



$\log N(\text{HI})=21.7$

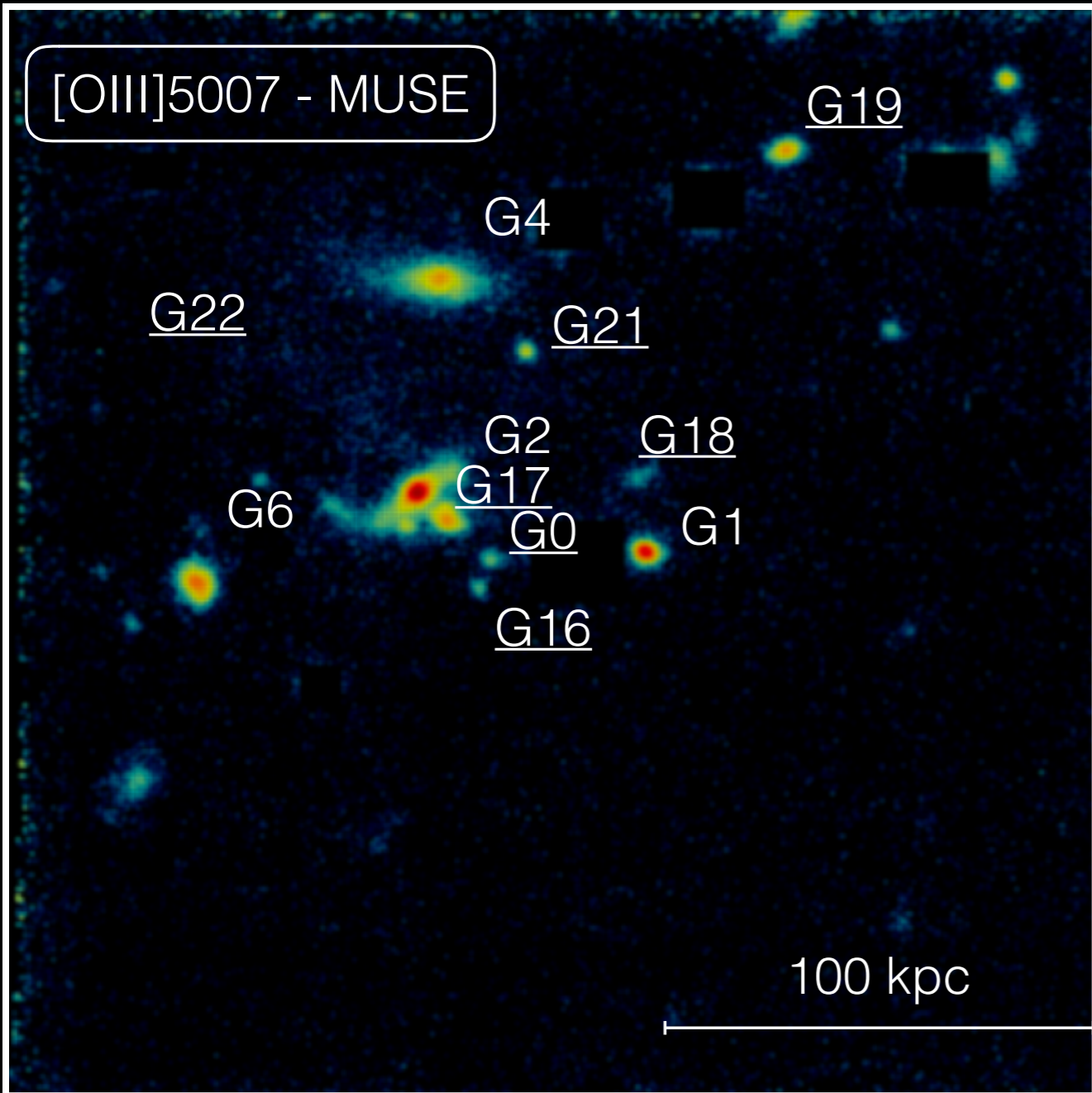
Kacprzak+10, Christensen+14

Absorber related to Small Group

[OIII]5007 - MUSE

G22 G4 G19
G6 G2 G21 G18
G17 G1 G0
G16

100 kpc



Large Molecular Gas Reservoirs

[OIII]5007 - MUSE

G19
G4
G22
G21
G2
G18
G17
G0
G1
G6
G16

100 kpc

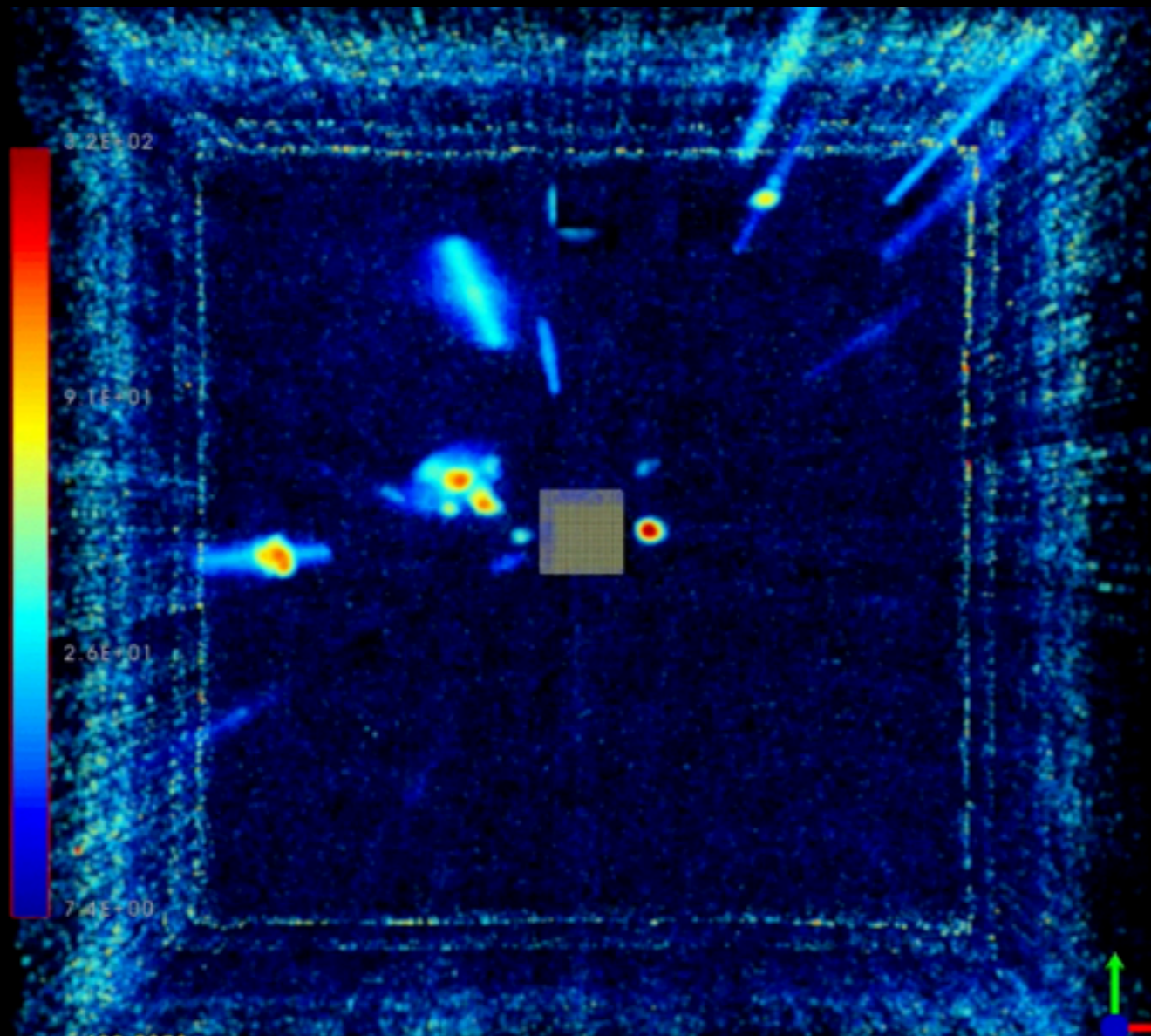
CO(1-0) - ALMA

G4
G2
G6

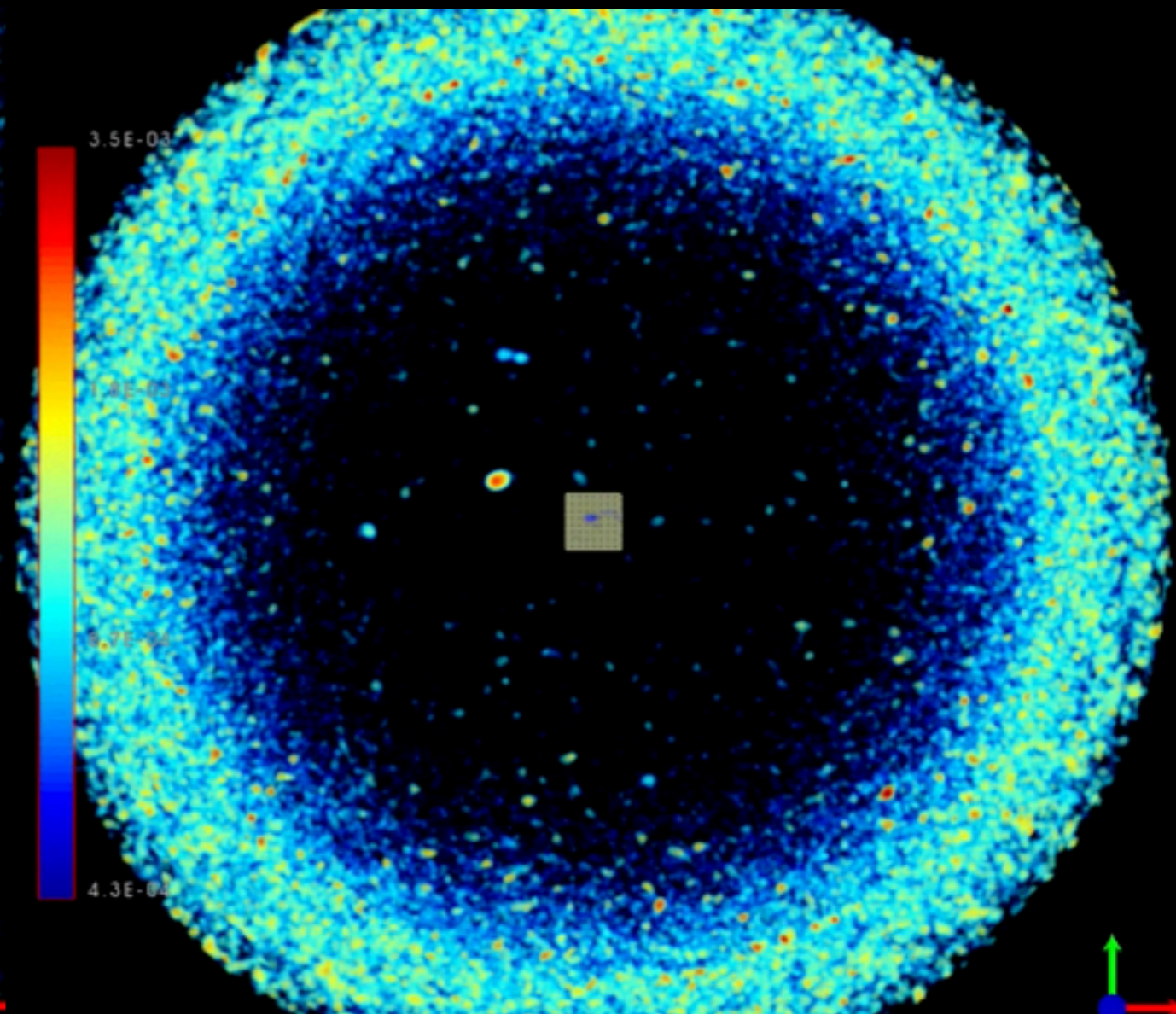
100 kpc

also Klitsch+18, Moller+18

Ionised Gas



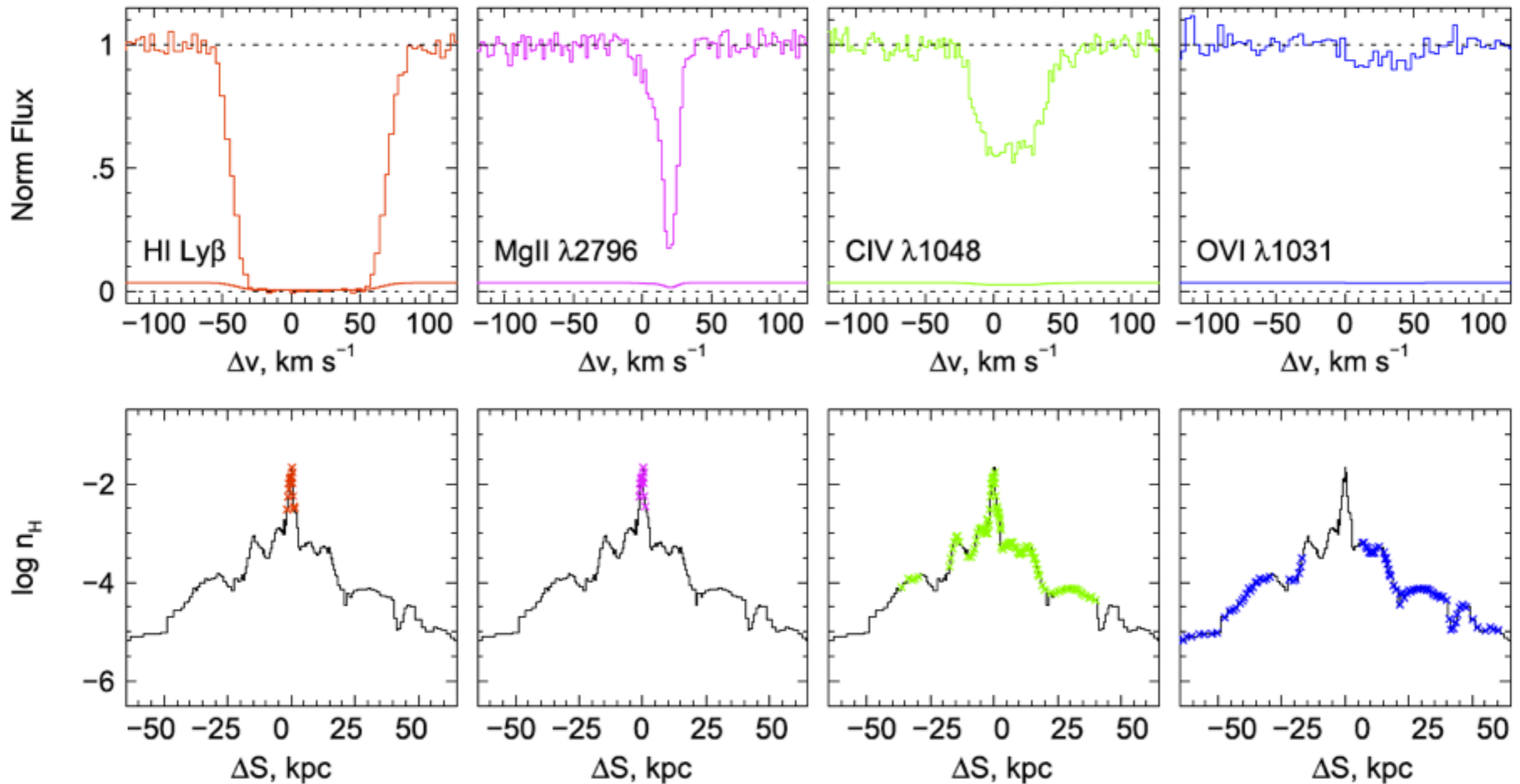
molecular gas



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Cold Gas Mixing

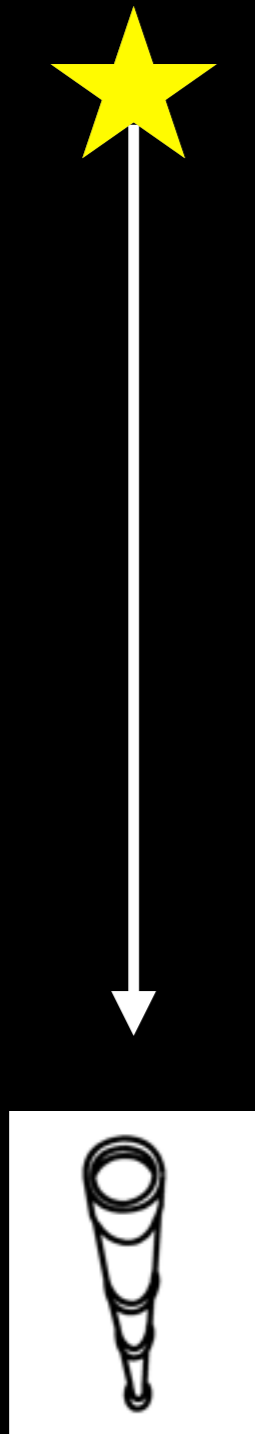


Churchill+15

Schaye+05,07; Gronke+17

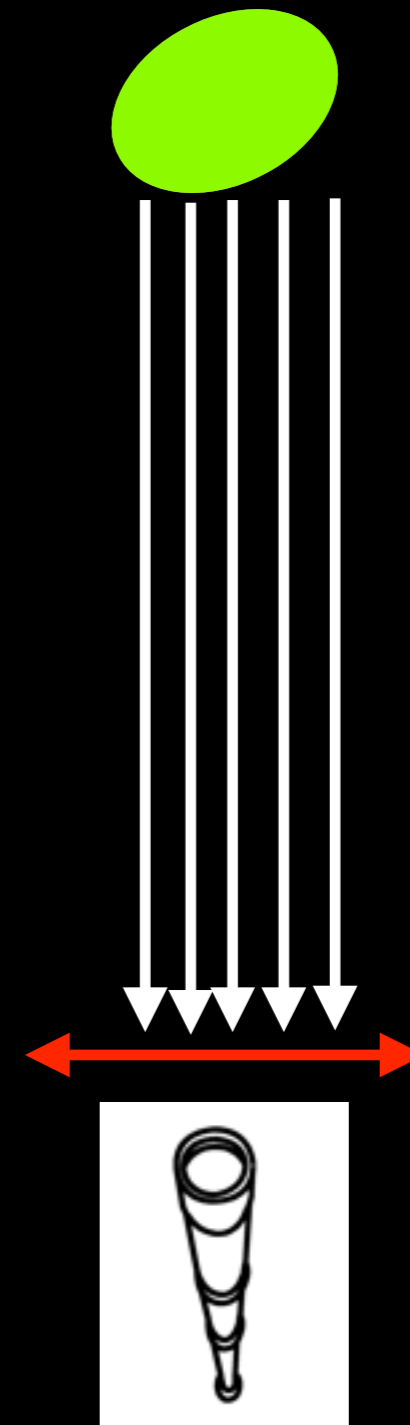
Beyond 1D along the Line-of-Sight

quasar (pt-like)



extended source

kpc-scale
in transverse
direction

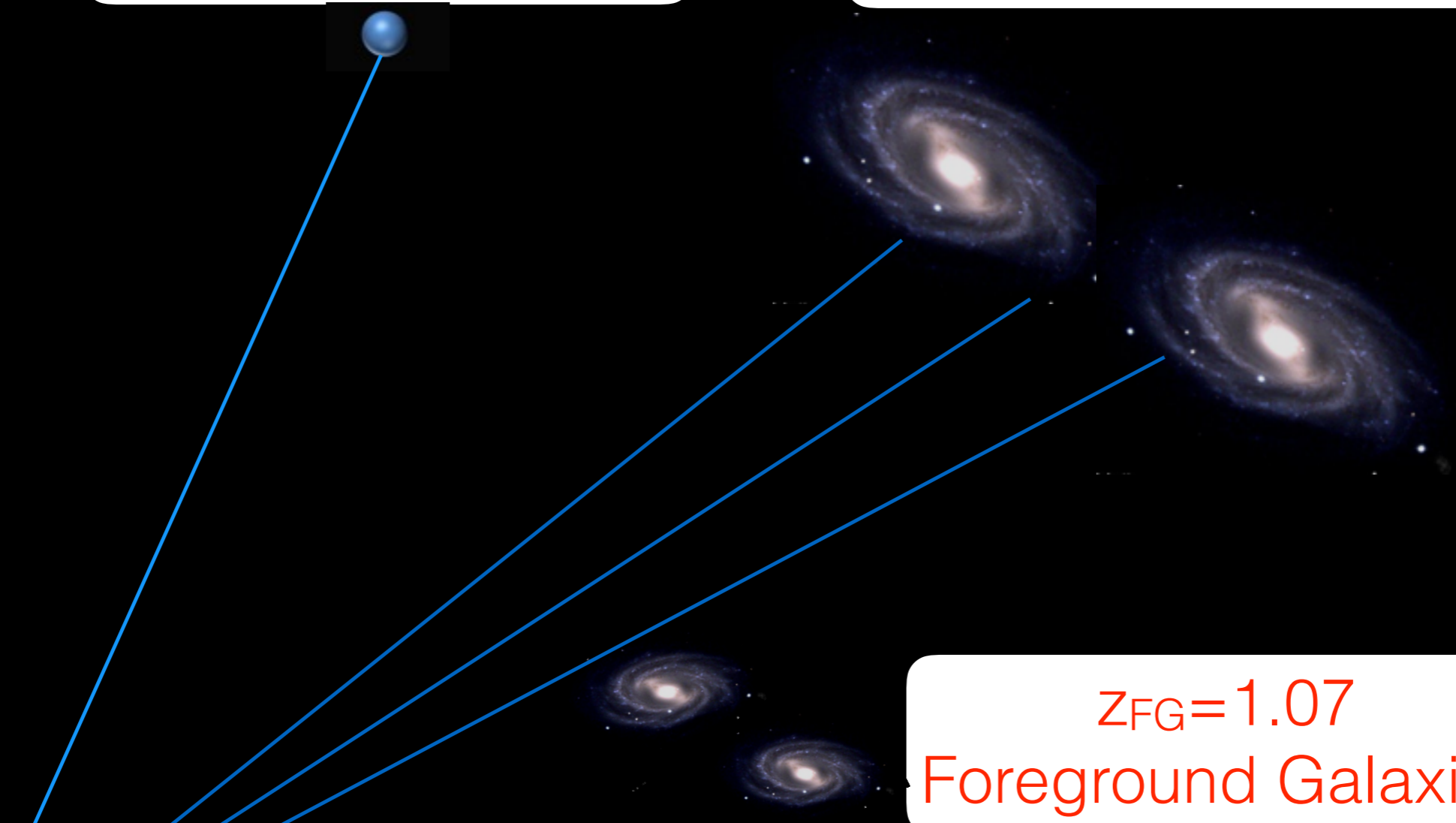


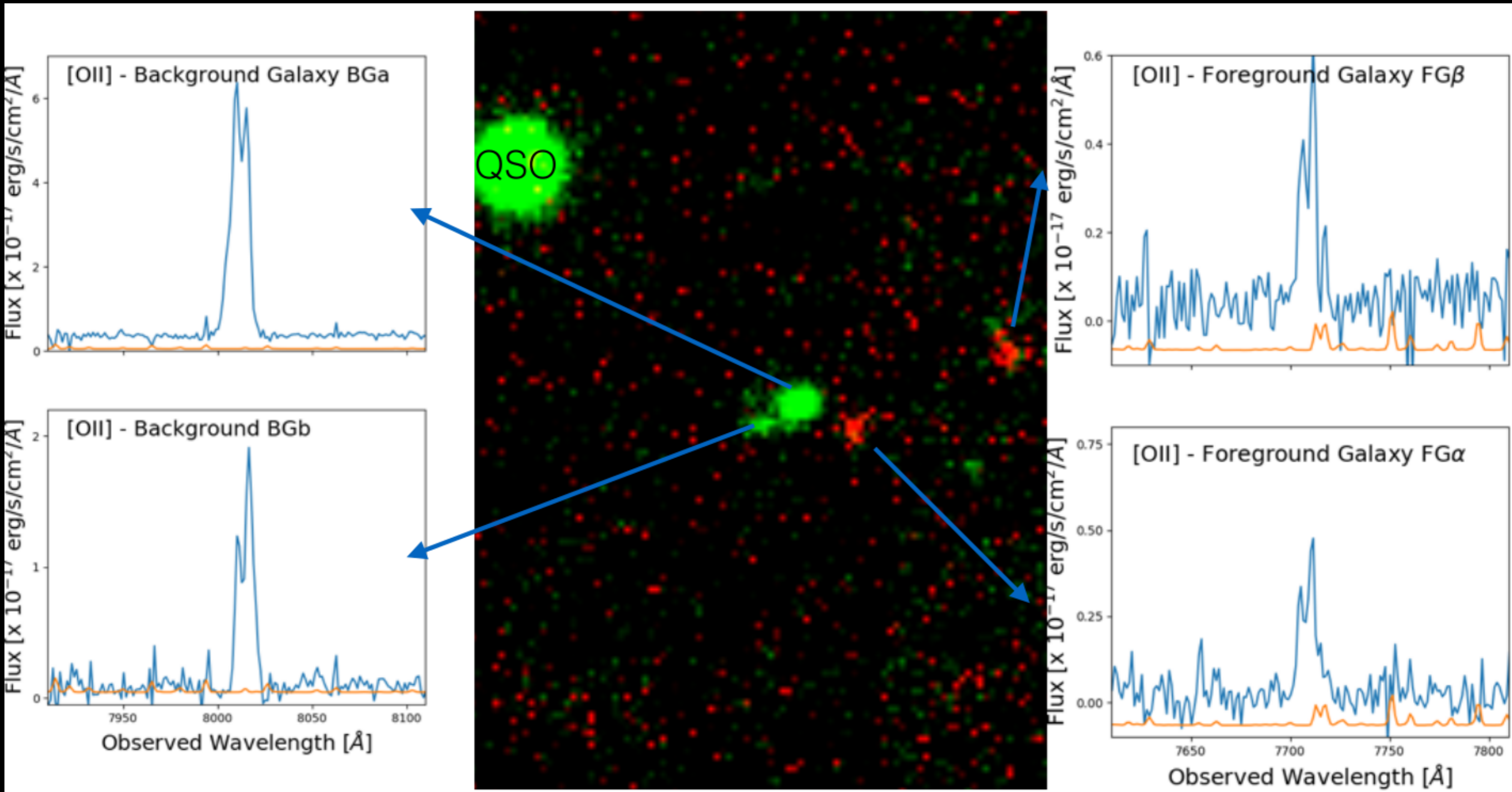
$z_{\text{QSO}}=3.33$
Background Quasar

$z_{\text{BG}}=1.15$
Background Galaxies

$z_{\text{FG}}=1.07$
Foreground Galaxies

$z_{\text{abs}}=1.07$
absorbers in
Background
Galaxy spectra



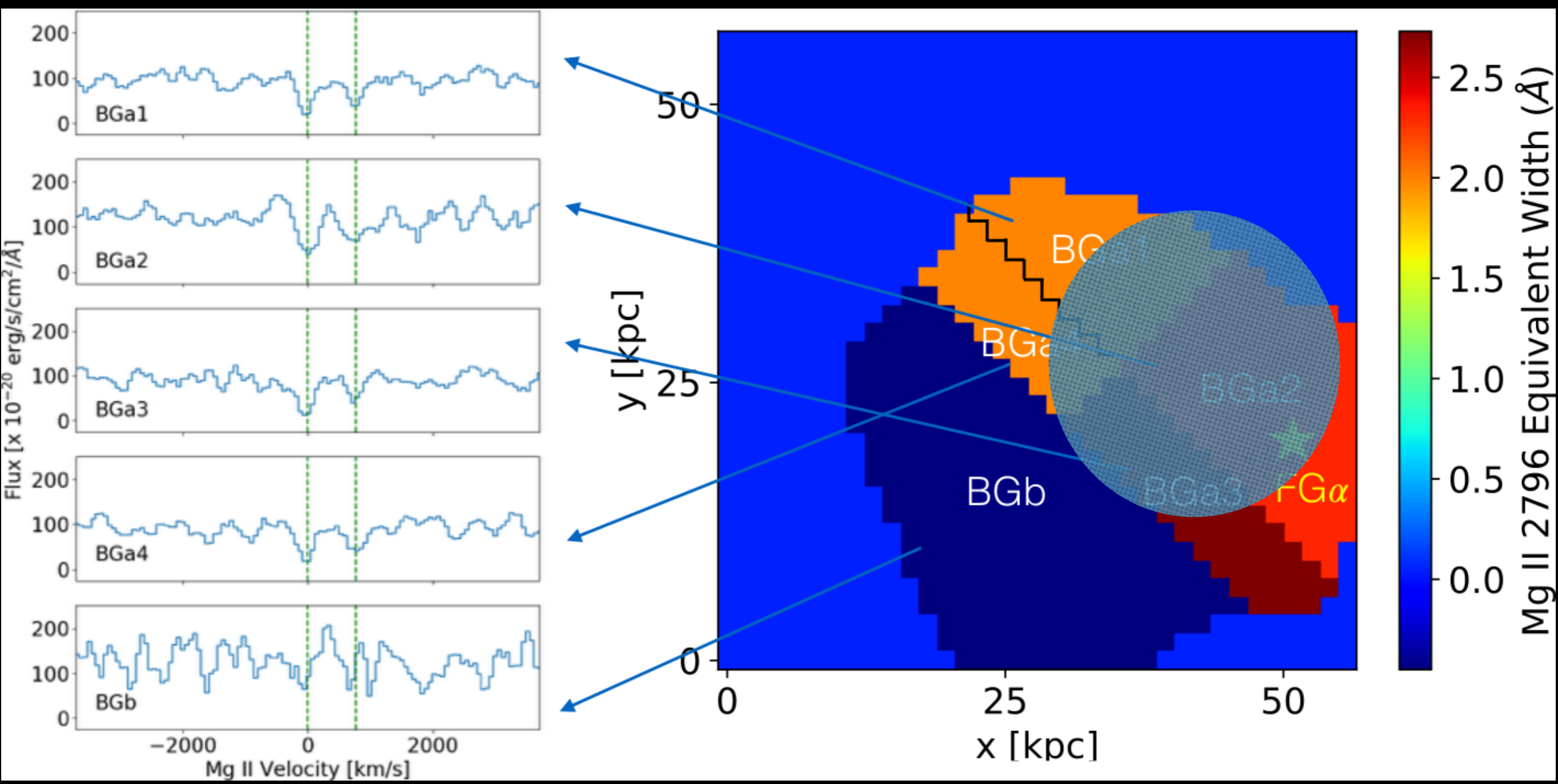


total area = 30 kpc^2

Spatially Resolved Metal Clouds

Cloud Gas Mass $< 2 \times 10^9 M_{\text{sun}}$

CP+18



Take Home Messages

extended source



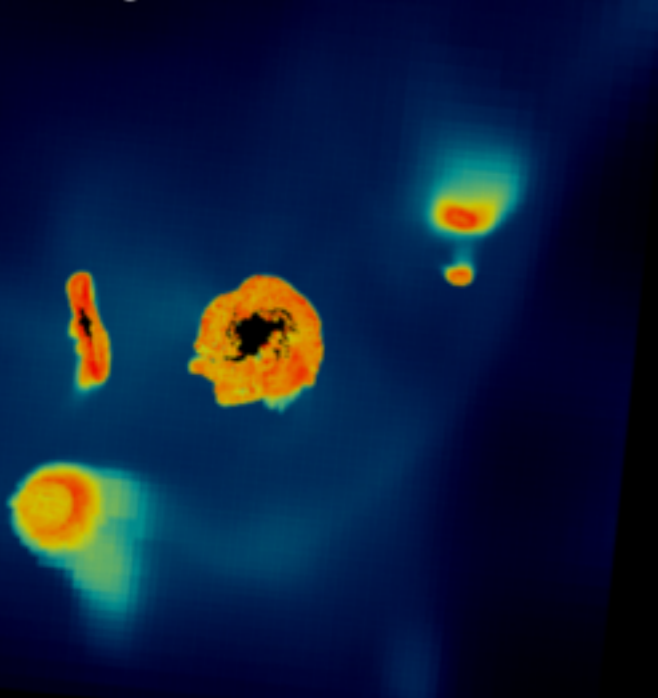
kpc-scale
in transverse
direction



Absorbers best tool to
date to characterise
low-surface
brightness tidal gas

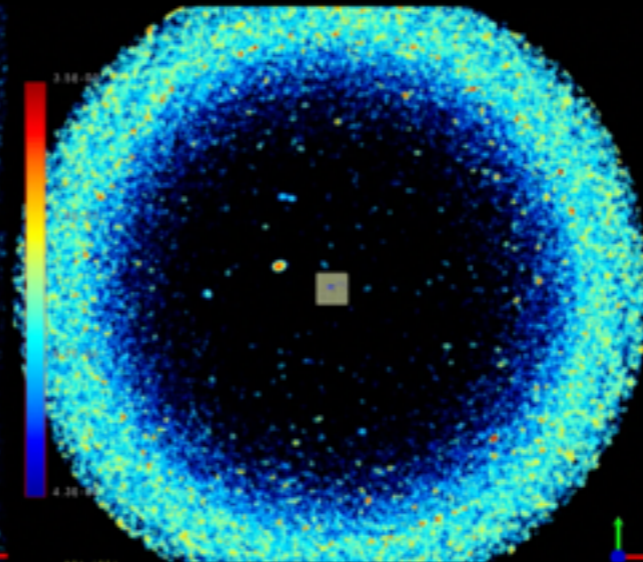
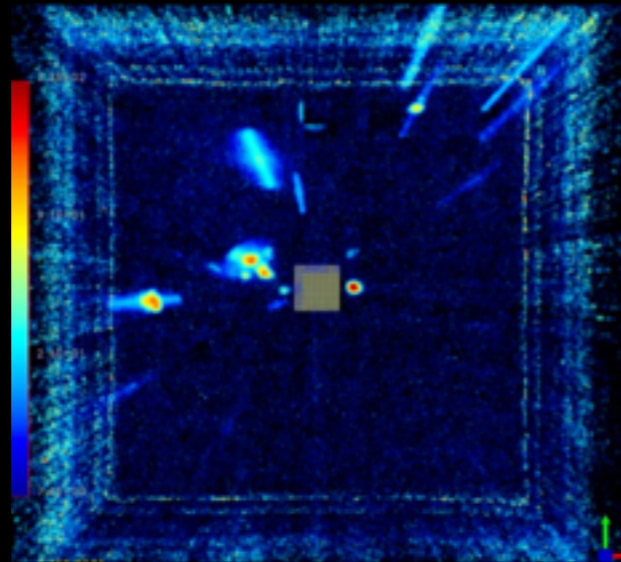
Ha Cosmological Zoom-in Simulations

8.2E-18
1.2E-19
1.9E-21
2.9E-23



Ionised Gas

molecular gas



Good efficiency
of the metal
mixing as traced
by cold gas

Absorbers
inefficiently
convert gas
into stars